Building a better non-uniform fast Fourier transform

ICERM 3/12/18

Alex Barnett (Center for Computational Biology, Flatiron Institute)

This work is collaboration with Jeremy Magland.

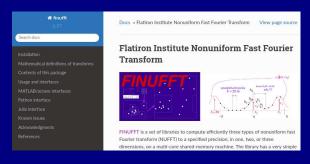
We benefited from much discussions and/or codes, including:
Leslie Greengard, Ludvig af Klinteberg, Zydrunas Gimbutas, Marina Spivak,

Joakim Andén, and David Stein

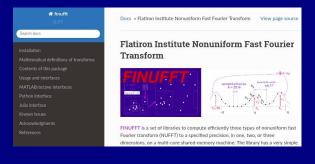




We are releasing https://github.com/ahbarnett/finufft

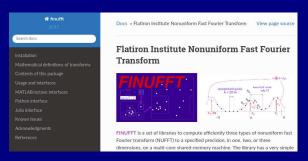


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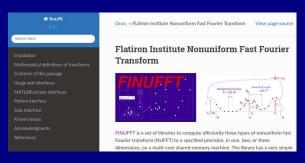
"Fourier analysis of non-uniformly spaced data at close to FFT speeds"

But...there already exist libraries?

eg NFFT from Chemnitz (Potts-Keiner-Kunis), NUFFT from NYU (Lee-Greengard)

Ours is faster in large-scale 2D and 3D settings & simpler to use

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Goals: show some math and engineering behind why, give applications...

... and explain how "Tex" Logan—one of the best bluegrass fiddle players in the country—is key to the story:



Task: eval.
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Vector
$$c$$
 of samples, $c_j = \frac{1}{N} f(\frac{2\pi j}{N})$

Easy to show its DFT is
$$f_k = \cdots + \hat{f}_{k-N} + \hat{f}_k + \hat{f}_{k+N} + \hat{f}_{k+2N} + \cdots$$

$$= \hat{f}_k \text{ desired } + \sum_{m \neq 0} \hat{f}_{k+mN} \text{ aliasing error due to discrete sampling}$$

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Key:
$$f = f = f$$
 smooth $\Leftrightarrow \hat{f}_n = f$ decays for $f = f$ large $\Leftrightarrow f$ aliasing error small $f = f$ bounded f

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Key:
$$f ext{ smooth } \Leftrightarrow \hat{f}_n ext{ decays for } |n| ext{ large } \Leftrightarrow ext{ aliasing error small}$$
 eg $(d/dx)^p f$ bounded $\Rightarrow \hat{f}_n = \mathcal{O}(1/|n|^p)$ p-th order convergence of error vs N

3) Estimate power spectrum of *non-periodic* signal on U grid

Contrast: what does NUFFT compute?

Inputs:
$$\{x_j\}_{j=1}^M$$
 NU (non-uniform) points in $[0,2\pi)$ $\{c_j\}_{j=1}^M$ complex strengths, N number of modes

Three flavors of task:

Type-1: NU to U, evaluates
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Type-2: U to NU, evaluates
$$c_j = \sum_{\substack{-N/2 \le k < N/2 \\ \text{Evaluate years}}} e^{ikx_j} f_k$$
, $j = 1, \dots, M$

Evaluate usual F series at arbitrary targets. Is adjoint (but not inverse!) of type-1

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Type-3: NU to NU, also needs NU output freqs $\{s_k\}_{k=1}^N$

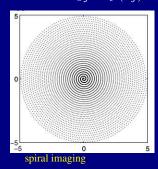
evaluates
$$f_k = \sum_{j=1}^M e^{is_k x_j} c_j \;, \quad k=1,\ldots,N$$
 general exponential sum

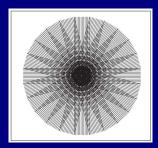
• For dimension d=2,3 (etc), replace kx_j by $\mathbf{k} \cdot \mathbf{x}_j = k_1x_j + k_2y_j$, etc

These tasks crop up a lot

• Magnetic resonance imaging (MRI).

f is unknown 2D image; seek its vector of values f on a U grid Given data $y_j = \hat{f}(\mathbf{k}_j)$ at NU set of Fourier pts \mathbf{k}_j :





PROPELLER

Evaluating the forward model, ie eval. $\mathbf{y} = A\mathbf{f}$, is a 2D type-2 NUFFT. Reconstruct \mathbf{f} by iteratively solving this (rect, ill-cond) linear system (eg Fessler) • Cryo electron microscopy (EM) algorithms (how we got into this)

We need to go between U voxel grid and NU spherical 3D **k**-space reps.



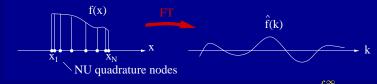
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• computing actual Fourier *transforms* of non-smooth $f(\mathbf{x})$ accurately:



Apply a good quadrature rule (eg Gauss) to the Fourier integral $\hat{f}(k) := \int_{-\infty}^{\infty} e^{ikx} f(x) dx$

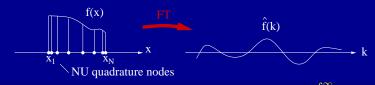
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- Coherent diffraction imaging again given Fourier data on NU pts (Ewald spheres)
- PDEs: interpolation from U to NU coords grids, applying heat kernels
- Making PDEs, mol. dyn, spatially periodic ("particle-mesh Ewald")
- Given large # of point masses (eg stars), what is Fourier spectrum?

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What is error? High freqs |k| = N/2 are the worst: relative error thus

$$e^{i\frac{N}{2}x_j} - e^{i\frac{N}{2}\tilde{x}_j} = \mathcal{O}(Nh) = \mathcal{O}(N/N_f)$$

- 1st-order convergent: eg error 10^{-1} needs $N_f \approx 10N$.
- in 3D needs $N_f{}^3 pprox 10^3 N^3$ 1000× slower than plain FFT, for 1-digit accuracy! Terrible!

And yet the idea of dumping onto fine grid is actually good...

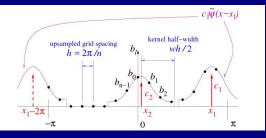
But need much more rapid convergence!

1D type-1 NUFFT algorithm

Three steps: Set up "not-as-fine" grid on $[0,2\pi),\ N_f=\sigma N,\$ upsampling $\sigma\approx 2$

Pick a $spreading \ kernel \ \psi(x)$ support must be only a few h wide

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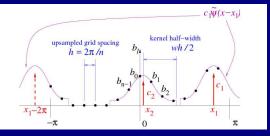


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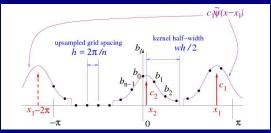
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- b) Do size- N_f FFT to get \hat{b}_k
- c) Correct for spreading: $\tilde{f}_k = \frac{1}{\hat{\psi}(k)}\hat{b}_k$, for $-N/2 \le k < N/2$

Why? since you convolved sum of point masses $\sum_{j=1}^{M} c_j \delta(x - x_j)$ with $\psi(x)$, undo by deconvolving: dividing by kernel in Fourier domain

Type-2 similar; type-3 needs more upsampling (by σ^2 not σ)

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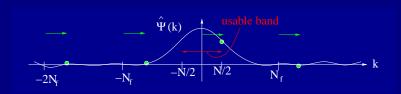
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A) and B) are conflicting requirements :(

Rigorous error analysis:
$$|\tilde{f}_k - f_k| \leq \epsilon \|\boldsymbol{c}\|_1$$
 where
$$\epsilon = \max_{|k| \leq N/2, x \in \mathbb{R}} \frac{1}{|\hat{\psi}(k)|} \Big| \sum_{m \neq 0} \hat{\psi}(k + mN_f) e^{i(k + mN_f)x} \Big|$$



want $\hat{\psi}$ large in |k| < N/2, small for $|k| > N_f - N/2$

(Partial) history of the NUFFT

- Interpolation of F series to NU pts, astrophysical (Boyd '80s, Press–Rybicki '89)
- Gaussian kernel case $\psi(x)=e^{-\alpha x^2}$ (Dutt-Rokhlin '93, Elbel-Steidl '98) rigorous proof of exponential convergence vs w, ie # digits = $\log_{10}(1/\epsilon) \approx 0.5w$
- Realization there's a close-to-optimal kernel ("Kaiser–Bessel") (Jackson '91) nearly twice the convergence rate: $\log_{10}(1/\epsilon) \approx 0.9w$ rigorous analysis (Fourmont '99, Fessler '02, Potts–Kunis...'02)
- Fast gridding for Gaussian case by cutting e^x evals (Greengard-Lee '04)
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But what on earth is Kaiser–Bessel?

Turns out requirements A) and B) v. close to those for good window funcs

Recall a window func, designed to make non-periodic signal pretend to be periodic...

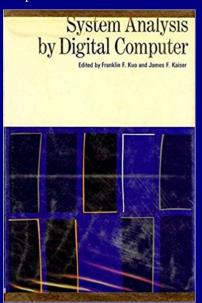
Story of the "Kaiser-Bessel" kernel

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(had to buy 2nd-hand) Let's open...

01

$$H_1^*(z) = \frac{1}{2} \sum_{n=1}^{N} [b_n^W(nT)][z^n - z^{-n}], \text{ m odd}$$
 (7.16)

where N is the greatest integer in (7/T).

An especially simple weighting function which may be used is Hamming's [13, p. 95-99] window function defined by

$$w_{h}(t) = \begin{cases} 0.54 + 0.46 \cos(\pi t/\tau), |t| < \tau \\ 0, |t| > \tau \end{cases}$$
 (7.17)

For this function 99.96% of its energy lies in the band $\mid \omega \mid \leq 2\pi/\tau$ with the peak amplitude of the side lobes of $W_h(j\omega)$ being less than 1% of the peak.

"Other specific window functions [38, 39] may be used to advantage. One especially flexible family of weighting functions with nearly optimum characteristics is given by the Fourier cosine transform pair [40]

$$w(t) = \begin{cases} \frac{\mathbb{I}_{0} \left[\omega_{0} \sqrt{\tau^{2} - t^{2}} \right]}{\mathbb{I}_{0}(\omega_{0}\tau)} & |t| < \tau \\ 0 & |t| > \tau \end{cases}$$
(7.18)

$$W(j_{\mathcal{B}}) = \frac{2}{T_{\mathcal{O}}(\omega_{\mathcal{A}}\tau)} \frac{\sin \left[\tau \sqrt{\omega^{1} - \omega_{\mathcal{A}}^{2}}\right]}{\sqrt{\omega^{2} - \omega_{\mathcal{A}}^{2}}}$$
(7.19)

where L_i is the modified Bessel function of the first kind and order zero. By varying the product ω_{τ} , the energy in the central lobe and the amplitudes of the side lobes can be changed. The usual range on values of ω_{τ} is $4 < \omega_{\tau} \tau < 9$ corresponding to a range of side lobe peak heights of 3.1% down to 0.047%. Figure 7.7 shows w(t) and W(j ω) for $\omega_{\tau} \tau = 6.0$ and 8.5

The flexibility of this set of window functions is now linearized. Setting $\omega_{\tau} = \pi_{\tau}/3 = 5.4441$ galacen the first zero of W(je) at 2π on a normalized scale, the same location as that of the cosine transform of the Hamming window [13, p. 95-99]. The closeness of the resulting windows is immediately apparent. The Hamming window, while having a slightly lower peak amplitude on the first two side lobes, continues to oscillate approximately sinusoidally with a slowly diminishing amplitude. On the other hand the L_{τ} -sink window with $\omega_{\tau} = \sqrt{3} \pi$ has a slightly greater magnitude for the first two side lobes but the amplitude of the sinusoidal oscillations of the window tails diminishes much more rapidly. The Hamming window has all but 0.07% of its energy in the main lobe while the L_{τ} -sin window with $\omega_{\tau} = \sqrt{3} \pi$ has all but 0.02% of its energy in the main lobe while the L_{τ} -sin window with $\omega_{\tau} = \sqrt{3} \pi$ has all but 0.02% of its energy in the main

As a second example, the Blackman window, his "not very serious proposal" [13, p. 98], has its first zero at 3 π and corresponds to the I_{σ} -sinh window with $\omega_{\pi}\tau=2\sqrt{\pi}\tau=8.885$. The nearly equivalent value of $\omega_{\pi}\tau=8.50$ can be used for comparison. This function has its first zero at $2\pi(1.44226)$ rather than at 3π . Again the closeness of these two window functions can be observed by comparing our Fig. 7. To with

[†] This family of window functions was "discovered" by Kaiser in 1962 following a discussion with B. F. Logan of the Bell Telephone Laboratories.

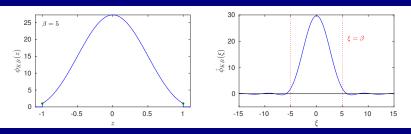
Kaiser-Bessel Fourier transform pair

The truncated kernel

$$\phi_{KB}(z) \ := \ \left\{ \begin{array}{ll} I_0(\beta\sqrt{1-z^2}), & |z| \leq 1 \\ 0, & \text{otherwise} \end{array} \right. \quad \text{we scale } z := \frac{2x}{wh}$$

is the FT of

$$\hat{\phi}_{KB}(\xi) = 2 \frac{\sinh \sqrt{\beta^2 - \xi^2}}{\sqrt{\beta^2 - \xi^2}}$$



Still unknown to Gradshteyn-Ryzhik, Bateman, Prudnikov, Wolfram ®, Maple ® . . .

Over to James Kaiser, Bell Labs (interviewed in '97)

About 1960-'61, Henry Pollak, who was department head in the math research area at Bell Labs, and two of his staff, Henry Landau, and Dave Slepian, solved the problem of finding that set of functions that had maximum energy in the main lobe consistent with certain roll-offs in the side lobes.

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Back to Kaiser...

So one day I went in Ben's office and his chalkboard was just filled with equations. Way down in the left-hand corner of Ben's chalkboard was this transform pair, the I_0 -sinh transform pair. I didn't know what I_0 was, I said, "Ben, what's I_0 ?" He came back with "Oh, that's the modified Bessel function of the first kind and order zero." I said, "Thanks a lot, Ben, but what is that?" He said, "You know, it's just a basic Bessel function but with purely imaginary argument." So I copied down the transform pair and went back to my office.

Back to Kaiser...

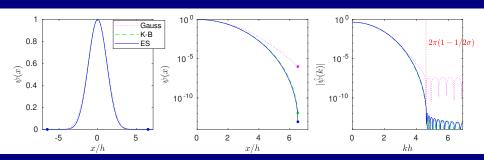
So one day I went in Ben's office and his chalkboard was just filled with equations. Way down in the left-hand corner of Ben's chalkboard was this transform pair, the I_0 -sinh transform pair. I didn't know what I_0 was, I said, "Ben, what's I_0 ?" He came back with "Oh, that's the modified Bessel function of the first kind and order zero." I said, "Thanks a lot, Ben, but what is that?" He said, "You know, it's just a basic Bessel function but with purely imaginary argument." So I copied down the transform pair and went back to my office.

I wrote a program ... got the data back and when I compared the I_0 function to the prolate, I said, "What's going on here? They look almost identical!" The answers were within about a tenth of a percent of one another. One program [PSWF] required 600 lines of code and the other ten or twelve lines of code!

P.S. we now have a kernel needing < 1 line of code ...

Compare the kernels

Plot the kernels for support of w = 13 fine grid points:



- very hard to distinguish on linear plot! Decays differ on log plot
- Kaiser-Bessel: tail of FT is at 10^{-12}
- best truncated Gaussian has tail only at 10^{-7}
- "ES" is our new kernel; v. close to KB

Our new ES ("exp of sqrt") kernel

$$\psi_{ES}(x) := e^{\beta\sqrt{1-z^2}}, \quad z := 2x/wh \in [-1,1],$$
 zero otherwise.

(found via numerical tinkering: simplifying the I_0)

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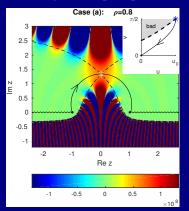
- its Fourier transform $\hat{\psi}_{ES}$ has no known formula
- 1) Numerical consequence: use quadrature on FT to eval. $\frac{1}{\hat{\psi}_{ES}}$ for step c)
- 2) Analytic consequence: one has to work with the FT integral directly...

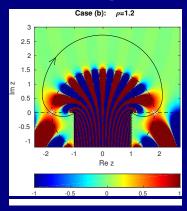
We prove essentially
$$\epsilon = \mathcal{O}(\sqrt{w}e^{-\pi w\sqrt{1-1/\sigma}})$$
 as kernel width $w \to \infty$

- same exponential convergence rate as Kaiser–Bessel, and as PSWF (Fuchs '64)
- consequence: $w \approx 7$ gives accuracy $\epsilon = 10^{-6}$, $w \approx 13$ gets $\epsilon = 10^{-12}$.
- However, evaluation now requires only one sqrt, one e^x , couple of mults.
- proof is 8 pages: contour integrals split into parts, sums into various parts, bounding the conditionally-convergent tail sum...

One proof ingredient

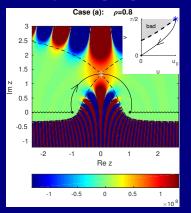
Asymptotics (in β) of the Fourier transform $\hat{\psi}(\rho\beta) = \int_{-1}^{1} e^{\beta(\sqrt{1-z^2}-i\rho z)} dz$ via deforming to complex plane, steepest descent (saddle pts)

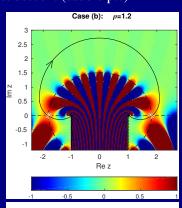




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Summary: new approx. FT pair

K-B & PSWF also may be interpreted this way

"exp(semicircle) $\stackrel{\text{FT}}{\longleftrightarrow}$ exp(semicircle) + exponentially small tail"

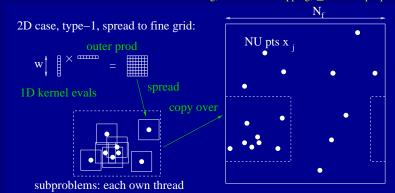
Implementation aspects

- Type 1,2,3 for dimensions d = 1, 2, 3: nine routines
- C++/OpenMP/SIMD, shared mem, calls FFTW. Apache license
- Wrappers to C, Fortran, MATLAB, octave, python, julia

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- Wrappers to C, Fortran, MATLAB, octave, python, julia
- Cache-aware multithreaded spreading:

Type-2 easy: parallelize over *bin-sorted* NU pts no collisons reading from U blocks Type-1 not so: writes collide load-balancing, slow index-wrapping, $< 10^4$ NU pts per subprob:



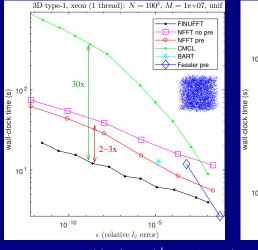
Performance: 3D Type-1 (the most dramatic)

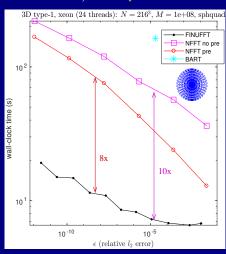
Compare FINUFFT to • CMCL NUFFT (single-threaded, Gaussian kernel)

• NFFT (multi-threaded, "backwards" Kaiser–Bessel) ie they eval. sinch $\sqrt{1-x^2}$

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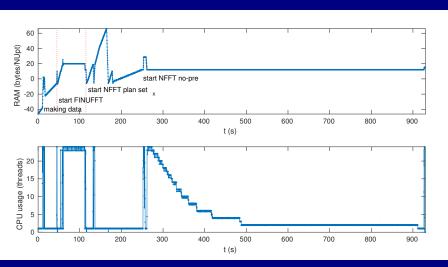




- all scale as $\mathcal{O}(M|\log\epsilon|^d+N\log N)$; it's about prefactors and RAM usage
- at $M = 10^8$: we need only 2 GB, vs NFFT pre needs 60 GB at high acc.

3D Type-1: RAM & CPU usage for non-uniform density

We use all threads efficiently, vs NFFT assigns threads to fixed x-slices:



Conclusions

NUFFT is a key tool with many scientific computing applications

We speed up and simplify the NUFFT using...

- mathematics: creation and rigorous analysis of new kernel func ψ
- no analytic $\hat{\psi}$ need be known: instead use numerical quadrature
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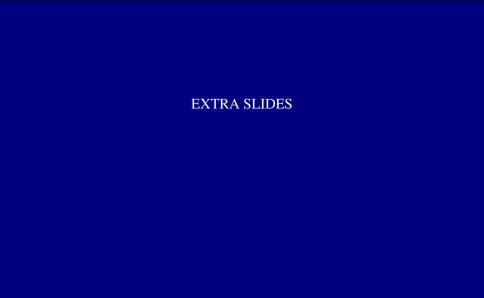
Result: FINUFFT (Flatiron Institute Non-Uniform Fast Fourier Transform)

```
https://github.com/ahbarnett/finufft
```

fast, simple to install and use. Send me bug reports & feature req's

Future:

- GPU spreader (build upon promising work of: Kunis–Kunis '12, Ou '17)
- math: "why" are PSWF and K–B so close to $e^{\beta\sqrt{1-z^2}}$? no, it's not WKB...



Vector intrinsics accelerate by up to $2\times$:

(Ludvig af Klinteberg)

```
out[0] = 0.0; out[1] = 0.0;
#if defined(VECT) && !defined(SINGLE)
   m128d vec out = mm setzero pd();
#endif
 if (i1>=0 && i1+ns<=N1 && i2>=0 && i2+ns<=N2 && i3>=0 && i3+ns<=N3) {
    // no wrapping: avoid ptrs
    for (int dz=0; dz<ns; dz++) {
      BIGINT oz = N1*N2*(i3+dz):
                                      // offset due to z
      for (int dy=0; dy<ns; dy++) {
        BIGINT i = oz + N1*(i2+dv) + i1:
       FLT ker23 = ker2[dv]*ker3[dz]:
        for (int dx=0; dx<ns; dx++) {
         FLT k = ker1[dx]*ker23:
#if defined(VECT) && !defined(SINGLE)
          m128d vec k = mm set1 pd(k);
          m128d vec val = mm load pd(du+2*j);
          vec out = mm add pd(vec out, mm mul pd(vec k, vec val));
#else
          out[0] += du[2*j] * k;
          out[1] += du[2*i+1] * k:
#endif
                                     wraps somewhere: use ptr list (slower)
```

• exploit SSE, SSE2, AVX, etc, common to 99% of CPUs.